

## Definition of an energy density and a generalized state of total radiation

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If one is given a time direction  $\mathbf{u}$  then a convenient choice of the three tensors will permit one to define the energy density scalar that is associated with  $\mathbf{u}$ . We define a state of total radiation by imposing some conditions on the curvature tensor of  $V_4$  that generalize the ones that A. Lichnerowicz pointed out <sup>(1)</sup>.

**1.** Let  $V_4$  be the space-time manifold of general relativity endowed with the metric  $ds^2 = g_{\alpha\beta} dx^\alpha dx^\beta$  <sup>(2)</sup>, let  $x_0$  be a point of  $V_4$ , let  $T_{x_0}$  be the tangent vector space at  $x_0$ , and let  $T_{x_0}^{\wedge(2)}$  be the subspace of  $T_{x_0}^{\wedge(2)}$  of  $T_{x_0}^{\wedge}$  antisymmetric tensors of order 2. If one is given a basis  $\mathbf{e}_{(\alpha)}$  for then we will suppose that  $T_{x_0}^{\wedge(2)}$  is referred to the basis <sup>(3)</sup>:

$$\mathbf{E}_{(I)} = \mathbf{e}_{(\alpha)} \wedge \mathbf{e}_{(\beta)}.$$

One will then have:

$$G_{IJ} \equiv \mathbf{E}_{(I)} \cdot \mathbf{E}_{(J)} = \gamma_{\alpha\beta, \lambda\mu} = g_{\alpha\beta, \lambda\mu} - g_{\alpha\lambda, \beta\mu}.$$

It is easy to associate the curvature  $R_{\alpha\beta, \lambda\mu}$  that is defined at  $x_0$  with the tensors:

$$*R_{\alpha\beta, \lambda\mu} = \frac{1}{2} \eta_{\alpha\beta\gamma\delta} R^{\gamma\delta, \lambda\mu} \quad \text{and} \quad **R_{\alpha\beta, \lambda\mu} = \frac{1}{2} \eta_{\alpha\beta\gamma\delta} \eta_{\lambda\mu\nu\sigma} R^{\gamma\delta, \lambda\mu},$$

in which  $\eta_{\alpha\beta\gamma\delta}$  is the volume element form. Those three tensors can be considered to be tensors of order 2 in  $T_{x_0}^{\wedge(2)}$ . In that case, they will be denoted by  $H_{II}$ ,  $*H_{II}$ , and  $**H_{II}$ , respectively. In what follows, we shall also utilize the three scalars:

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<sup>(1)</sup> A. LICHNEROWICZ, C. R. Acad. Sci. Paris **246** (1958), pp. 893.

<sup>(2)</sup>  $\alpha, \beta, \dots = 0, 1, 2, 3; i, j, \dots = 1, 2, 3$ .

<sup>(3)</sup> The correspondence between indices  $\alpha$  and I conforms to the substitution:

$$\begin{pmatrix} 23 & 31 & 12 & 10 & 20 & 30 \\ 1 & 2 & 3 & 4 & 5 & 6 \end{pmatrix}.$$

$$A = \frac{1}{2} H_{\mathbb{I}\mathbb{J}} H^{\mathbb{I}\mathbb{J}} = \frac{1}{2} **H_{\mathbb{I}\mathbb{J}} **H^{\mathbb{I}\mathbb{J}} = -\frac{1}{2} *H_{\mathbb{I}\mathbb{J}} *H^{\mathbb{I}\mathbb{J}},$$

$$B = \frac{1}{2} H_{\mathbb{I}\mathbb{J}} *H^{\mathbb{I}\mathbb{J}} = -\frac{1}{2} *H_{\mathbb{I}\mathbb{J}} **H^{\mathbb{I}\mathbb{J}},$$

$$C = \frac{1}{2} H_{\mathbb{I}\mathbb{J}} **H^{\mathbb{I}\mathbb{J}}.$$

2. If one is given a vector  $u^\alpha$  of square + 1 at a point  $x_0$  then consider the three tensors:

$$Y_{\beta\mu} = R_{\alpha\beta, \lambda\mu} u^\alpha u^\lambda, \quad X_{\beta\mu} = **R_{\alpha\beta, \lambda\mu} u^\alpha u^\lambda, \quad Z_{\beta\mu} = - *R_{\alpha\beta, \lambda\mu} u^\alpha u^\lambda.$$

With respect to any orthonormal frame such that  $\mathbf{e}_{(0)} = \mathbf{u}$ , one will have:

$$A = \frac{1}{2} (X_{\beta\mu} X^{\beta\mu} + Y_{\beta\mu} Y^{\beta\mu} - Z_{\beta\mu} Z^{\beta\mu}), \quad B = (X_{\beta\mu} - Y_{\beta\mu}) Z^{\beta\mu}, \quad C = X_{\beta\mu} Y^{\beta\mu} + Z_{\beta\mu} Z^{\beta\mu}.$$

The spatio-temporal square of each of these three tensors is positive or zero; they will be annulled only if the corresponding tensor is zero. If all three are zero then the curvature tensor itself will be zero. Consider the scalar:

$$V = \frac{1}{2} (X_{\beta\mu} X^{\beta\mu} + Y_{\beta\mu} Y^{\beta\mu} + 2 Z_{\beta\mu} Z^{\beta\mu}).$$

This scalar is strictly positive unless  $R_{\alpha\beta, \lambda\mu}$  is zero. We call it the *energy density that is associated with the time direction  $\mathbf{u}$* <sup>(4)</sup>.

3. We propose to say that the point  $x_0$  presents a generalized total radiation state if the following hypotheses are satisfied:

$\mathcal{H}_1$ . There exists an isotropic vector  $l^\alpha$  such that:

$$R_{\alpha\beta, \lambda\mu} l^\alpha l^\beta = 0, \quad *R_{\alpha\beta, \lambda\mu} l^\alpha l^\beta = 0, \quad **R_{\alpha\beta, \lambda\mu} l^\alpha l^\beta = 0.$$

$\mathcal{H}_2$ . There exists a vector  $u^\alpha$  of square + 1 such that:

$$R_{\alpha\beta, \lambda\mu} u^\alpha u^\lambda l^\beta = 0, \quad **R_{\alpha\beta, \lambda\mu} u^\alpha u^\lambda l^\beta = 0.$$

If that were true then one would find an orthonormal frame  $\mathbf{e}_{(\alpha)}$  such that  $\mathbf{e}_{(0)} = \mathbf{u}$ ,  $\mathbf{e}_{(1)} + \mathbf{e}_{(2)} = \mathbf{l}$  with respect to which one will have:

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<sup>(4)</sup> For the Schwarzschild case, that definition will coincide with a result of SYNGE, Proc. Roy. Irish Acad. **58**, A4.

$$(H_{IJ}) = \begin{pmatrix} 0 & 0 & 0 & 0 & 0 & 0 \\ 0 & \alpha_2 & \beta & 0 & \beta & \sigma_2 \\ 0 & \beta & \alpha_2 & 0 & \sigma_3 & -\beta \\ 0 & 0 & 0 & 0 & 0 & 0 \\ 0 & \beta & \sigma_3 & 0 & \gamma_2 & -\beta \\ 0 & \sigma_2 & -\beta & 0 & -\beta & \gamma_3 \end{pmatrix}, \quad \sigma_2 = \frac{1}{2}(\gamma_3 + \alpha_2), \quad \sigma_3 = \frac{1}{2}(\gamma_2 + \alpha_3).$$

The proper values of  $(H_{IJ})$  with respect to  $(G_{IJ})$  are  $\rho_1 = \rho_4 = 0$ ,  $\rho_3 = \rho_5 = (1/2)(\alpha_3 - \gamma_2)$ ,  $\rho_2 = \rho_6 = (1/2)(\alpha_2 - \gamma_3)$ . The vectors in the 2-plane  $(E_{(1)}, E_{(4)})$ ,  $M \equiv E_{(3)} - E_{(5)}$  and  $L \equiv E_{(2)} + E_{(6)}$  are proper vectors that correspond to  $\rho_1$ ,  $\rho_3$ , and  $\rho_2$ , resp.  $H_{IJ}$  admits the reduction:

$$H_{IJ} = \beta(L_I M_J + L_J M_I) + \frac{1}{2} \alpha_2 (E_{(2)I} L_J + E_{(2)J} L_I) + \frac{1}{2} \alpha_3 (E_{(3)I} L_J + E_{(3)J} L_I) \\ - \frac{1}{2} \gamma_2 (E_{(5)I} L_J + E_{(5)J} L_I) + \frac{1}{2} \gamma_3 (E_{(6)I} L_J + E_{(6)J} L_I).$$

Similarly, the proper values of  $(R_{\alpha\beta})$  with respect to  $(g_{\alpha\beta})$  are  $s_0 = s_1 = R/4$ ,  $s_2 = 2\rho_3$ ,  $s_3 = 2\rho_2$  ( $s_0 + s_3 = R/2$ ).  $l$ ,  $e_{(2)}$ , and  $e_{(3)}$  are proper vectors that correspond to  $s_0$ ,  $s_2$ , and  $s_3$ , resp.  $S_{\alpha\beta} = R_{\alpha\beta} - (1/2)R g_{\alpha\beta}$  admits the reduction:

$$S_{\alpha\beta} = -\frac{1}{2}(\alpha_2 + \alpha_3 + \gamma_2 + \gamma_3) l_\alpha l_\beta - \frac{1}{2}(s_2 + s_3)(l_{(0)\alpha} l_{(0)\beta} - l_{(1)\alpha} l_{(1)\beta}) \\ + s_3 l_{(2)\alpha} l_{(2)\beta} + s_2 l_{(3)\alpha} l_{(3)\beta}.$$

In addition, one will have:

$$A = (\rho_2)^2 + (\rho_3)^2, \quad B = 0, \quad C = -2 \rho_2 \rho_3.$$

4. One will deduce from the relations:

$$e_{(2)}^\mu \nabla_\beta (*R_{\alpha}^{\beta}{}_{,\lambda\mu} l^\alpha l^\lambda) = 0, \quad e_{(3)}^\mu \nabla_\beta (*R_{\alpha}^{\beta}{}_{,\lambda\mu} l^\alpha l^\lambda) = 0$$

that:

$$s_2 e_{(3)\alpha} l^\beta \nabla_\beta l^\alpha = 0, \quad s_3 e_{(2)\alpha} l^\beta \nabla_\beta l^\alpha = 0.$$

If  $s_2 \neq 0$ ,  $s_3 \neq 0$  then it will result that the vector  $l^\beta \nabla_\beta l^\alpha$  is orthogonal to  $e_{(2)}$  and  $e_{(3)}$ . Therefore:

$$l^\beta \nabla_\beta l^\alpha = a l^\alpha.$$

If  $s_2 = s_3 = 0$  then  $R_{\alpha\beta, \lambda\mu}$  will satisfy the conditions:

$$R_{\alpha\beta, \lambda\mu} l^\alpha = 0, \quad *R_{\alpha\beta, \lambda\mu} l^\alpha = 0,$$

and the preceding result will remain true, from a theorem of A. Lichnerowicz <sup>(1)</sup>. In those two cases, we state: *The trajectories of the vector  $l$  that is associated with a*

*generalized state of total radiation that is defined in a domain are geodesics of the metric.* We point out that the case for which the theorem is not established is characterized by the relations  $A \neq 0, B = C = 0$ .

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